Small mounds and frost heave-like morphologies within Occator crater, Ceres^[1]

K. H. G. Hughson¹, B. E. Schmidt¹, H. G. Sizemore², J. E. C. Scully³, K. Duarte¹, V. Romero¹, P. M. Schenk⁴, D. L. Buczkowski⁵, D. A. Williams⁶, A. Nathues⁷, J. C. Castillo-Rogez³, C. A. Raymond³, C. T. Russell⁸.

Introduction

• Occator crater (Fig. 1) is one of the youngest and most prominent geologic structures on Ceres [2,3].

• Home to the largest of Ceres' carbonate deposits, Cerealia and Vinalia Faculae [4], Occator was imaged at \sim 3 m/px resolution by the Dawn spacecraft.

• Dawn imaged possible ice related features, including fractures, central pit craters, and many small mounds.

 Here we identify, categorize, and examine the small mounds and frost heave-like morphologies in Occator, and test the hypothesis that they may be generated by ground ice and hydrological processes.

Methods

 We employed 1:10,000 scale geologic mapping to identify all small mounds larger than 40 m in diameter.

 Mounds were classified based on common characteristics, such as: angularity, flatness, cliff forming layers, and summit depressions.

• We performed a geospatial analysis of the mounds using unit intersection and density-based clustering.

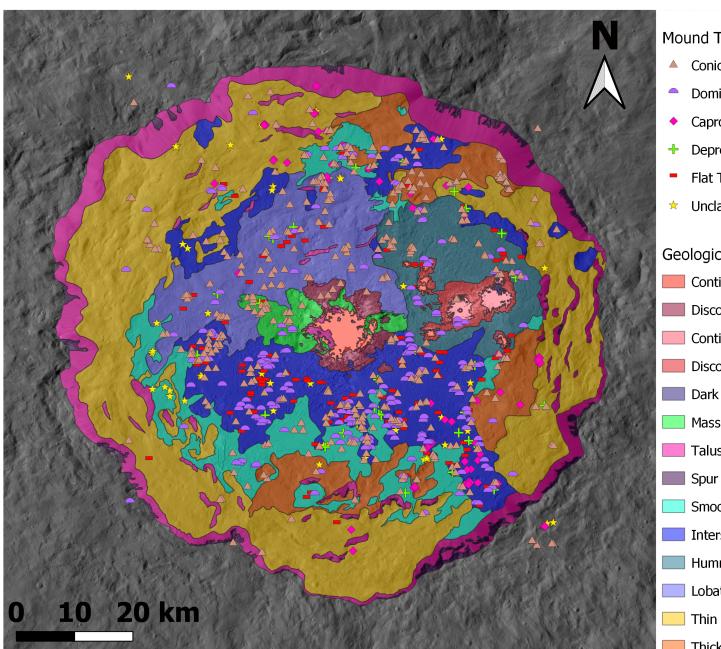


Fig. 1

- Depressed Sum

Results

• We identified 922 small mounds within Occator.

• The mounds are divided into six classifications (Fig. 2).

• Conical, domical, and depressed summit mounds are the most frost heave-like in character.

• 49% of mounds are found within Interspersed Lobate Material (ILM). This unit comprises 16% of Occator.

• The majority of statistically significant mound clusters are also found on ILM.

Discussion

 Occator has a large number mounds that appear independent of small craters and boulder fields.

 Many mounds have features such as fractures, pits, and conical-to-trapezoidal profiles, which are common among pingos (Fig. 3,4).

• The mounds have a strong affinity for ILM. This supports the hydrologic hypothesis as ILM is interpreted as an impact melt rich slurry [2].

• The morphological clustering of mounds on specific units may be indicative of increased hydraulic conductivity and historic melt reservoirs.

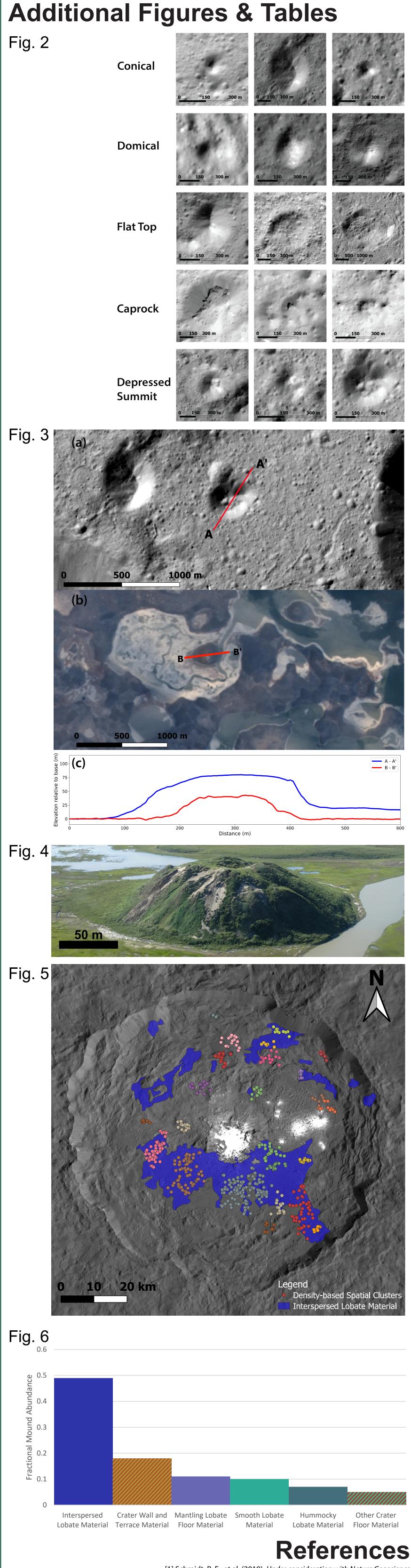


Frost Heaves and Pingos may exist on Ceres.

Their shapes and distribution could help characterize groundwater systems on frozen worlds.



1. School of Earth and Atmospheric Science, Georgia Institute of Technology, USA (khughson7@gatech.edu), 2. Planetary Science Institute, USA, 3. Jet Propulsion Laboratory, USA, 4. Lunar and Planetary Institute, USA, 5. Johns Hopkins Applied Physics Laboratory, USA, 6. Arizona State University, USA, 7. Max Planck Institute for Solar System Research, Germany, 8. University of California at Los Angeles, USA.



[1] Schmidt, B. E., et al. (2019). Under consideration with Nature Geoscience [2] Scully, J. E. C., et al. (2019). lcarus. https://doi.org/10.1016/j.icarus.2018.04.01 [3] Scully, J. E. C., et al. (2019). Under consideration with Nature Geoscience [4] De Sanctis, M. C., et al. (2016). Nature. doi:10.1038/nature18290 [5] Morrison, M. A. (2019). #betterposter. Retrieved from osf.io/ef53